

Characterization of silicon pore optics for the NewATHENA observatory in the PTB laboratory at BESSY II

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- 1. The **NewATHENA** observatory
- 2. The **PTB** laboratory at BESSY II
- 3. Mirror module assembly and characterization at the X-ray Parallel Beam Facility (XPBF 2.0)
- 4. Reflectance measurement on a mirror module at the Four Crystal Monochromator (FCM) beamline

NewATHENA facts and figures

PIB

ATHENA THE ADVANCED TELESCOPE FOR HIGH ENERGY ASTROPHYSICS

A mission addressing The Hot and Energetic Universe science theme

How did ordinary matter ASSEMBLE INTO THE LARGE SCALE STRUCTURES WE SEE TODAY?

How do black holes grow AND SHAPE THE UNIVERSE? Two instruments (detectors):

X-ray Integral Field Unit (X-IFU),

cryogenic imaging spectrometer, 0.3 to 10 keV

Wide Field Imager (WFI),

silicon active pixel sensor, 0.1 to 12 keV energy range

- Effective area of >1.1 m² at 1 keV
- Angular resolution of < 9 arcsec
- Field of view of > 40 arcmin



Silicon Pore Optics (SPO)

with 12 m focal length



European Space Agency





Silicon Pore Optics (SPO): mirror module production





cosine

measurement systems

38 plates are assembled to a stack by a stacking robot

2 stacks are assembled to an XOU (X-ray Optical Unit) for primary and secondary reflection



2 XOUs form a mirror module (MM)

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Braunschweig and Berlin







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PTB laboratory at BESSY II





Main parameters:

ring circumference	240 m
electron energy	1.7 GeV
max. ring current	300 mA, top-up
beamlines	50



- **B1** plane grating monochromator SX700 30 eV to 1800 eV **B2** four-crystal monochromator (FCM) 1.75 keV to 10 keV (0.7 nm to 0.1 nm) B3a undispersed bending magnet radiation **B5** B3b normal incidence monochromator **B6** 3 eV to 35 eV (400 nm to 35 nm) **B7** B3c deflected undispersed bending magnet radiation, EUV irradiation test station undispersed undulator radiation B4a **B**8 Compton backscattering
 - B4bplane grating monochromator (PGM) at undulator,
40 eV to 1900 eV (30 nm to 0.65 nm)
 - B4c deflected undispersed undulator radiation EUVL metrology test station
 - X-ray pencil beam facility (XPBF), astrophysics optics characterization
 - **X-ray parallel beam facility (XPBF 2.0),** astrophysics optics characterization
 - Tender X-ray microfocus beamline 1.5 keV to 10 keV (multilayer mirror monochromator & DCM)
 - 8 BAMline, double multilayer and double crystal monochromator (DCM & DMM), 8 keV to 60 keV

X-ray Parallel Beam Facility (XPBF 2.0)

- Cooperation between PTB and ESA for optics and detector characterization with synchrotron radiation since 1994 (BESSY I)
- X-ray Pencil Beam Facility (XPBF 1) in operation since 2005
- X-ray Parallel Beam Facility (XPBF 2.0) in operation since 2016, upgraded in 2020



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X-ray Parallel Beam Facility (XPBF 2.0)





Sample chamber of XPBF 2.0







Hexapod for sample (mirror module) positioning

- Travel ranges hor. plane: ± 60 mm, vertical ± 30 mm (+ 85 mm)
- Angular range: 15°, freely selectable pivot point
- Angular positioning accuracy: 0.7 arcsec using feedback from two electronic autocollimators





Detector 'tower'

- Vertical translation: 2.1 m
- Translation: in beam direction 1 m, perpendicular 100 mm
- Absolute detector position measurement with laser tracker
- Phosphor screen, tandem optics, CCD camera 2048 x 2048, 13.5 µm pixel size



Available beam at XPBF 2.0





Mirror Module (MM): combination of two XOUs

Assembled at XPBF 2.0 using 3 hexapods for positioning of the stacks

• X-rays are used for the alignment

•







cosine

measurement systems



Time evolution of **X**-ray **O**ptical **U**nit quality





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Four Crystal Monochromator (FCM) beamline





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X-ray fluorescence (XRF) scan on part of MM top layer





MM reflectance line scan









Photon energy scans through individual pores, double reflection Very stable beam position required!

- pores on 3 different plates (1, 9, 17)
- 2 off-center pores for plate 1 (± 20 mm)

Raster scans at fixed angle







Earlier measurements from January 2022 on a partly coated MM which has also been measured at the PGM beamline in the energy range from 0.2 to 1.8 keV **Double reflection peaks remains always at the same position**

Conclusion and outlook



- NewATHENA will be the largest X-ray telescope ever flown
- Mirror will be based on Silicon Pore Optics
- Mirror Modules are assembled and characterized at XPBF 2.0
- Energy-dependent reflectance has been measured on a Ir coated mirror module at the FCM beamline in the range from 1.75 to 10 keV
- Currently, a (partly modified) copy of XPBF 2.0 (called MINERVA) is being installed at the ALBA synchrotron radiation facility in Spain
- After final mission adoption, two additional beamlines will be installed in the PTB laboratory at BESSY II for the production of the about 500 mirror modules
- The launch of NewATHENA is currently scheduled for 2036, the budget for ESA is 1.3 G€

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